Astronomy 103 Final Exam May 12, 2014

Instructions:

No books, notes or calculator are allowed. You have 2 hours to complete the exam, and please do not turn to the next page until instructed to do so.

You may find the following information helpful:

- $1 \text{ AU} = 3 \times 10^8 \text{ km}$
- speed of light = 3×10^8 m/s
- Kepler's 3rd law:

$$a^3 = P^2$$

with the period P in years and semi-major axis a in AU.

• Newton's law of gravity:

$$F = \frac{GMm}{r^2}$$

• Peak wavelength and temperature of blackbody radiation:

$$\lambda = \frac{3\times 10^6}{T} \text{ nm}$$

with wavelength λ in nm (1 nm = 10^{-9} m) and temperature T in Kelvin.

- Conversion of mass into energy: $E = mc^2$
- Relationship between brightness *B* and distance *d*:

$$B_2 = B_1 \times \frac{d_1^2}{d_2^2}$$

• Relationship between luminosity L, temperature T and radius R of stars:

$$L = 4\pi\sigma T^4 R^2$$

where $4\pi\sigma$ are constants.

• Hubble's law:

$$v = H \times d$$

where v is velocity in km/s, d is distance in Mpc, and H is the Hubble constant, H = 70 km/s/Mpc.