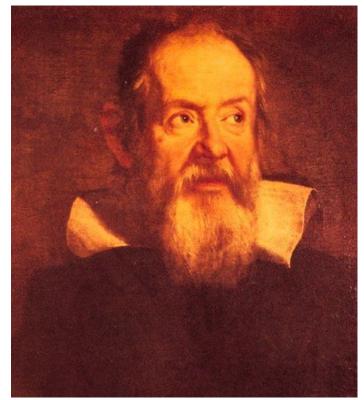
### Announcements

- Please finish reading Chapter 1
- Quiz 3 due tonight, 11.59pm
  - Problem Set 3 for practice
- Start reading Chapter 2
- Change in 3<sup>rd</sup> mid-term date: now Monday, Nov. 24<sup>th</sup> (was Nov 26<sup>th</sup>) i.e. Monday before Thanksgiving break

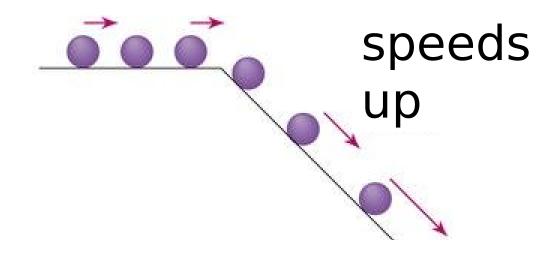
### Galileo's astronomical observations

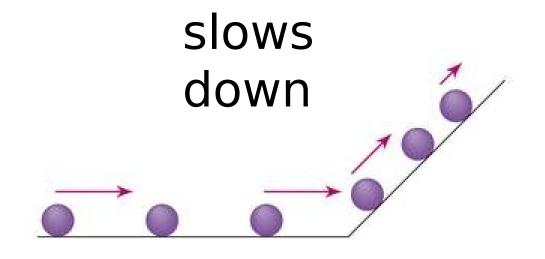
- Phases of Venus, agreeing with Copernican model
- Craters on the Moon
- 4 moons of Jupiter
- The Milky Way seen as a vast collection of stars
- Sunspots: motion showed that the Sun rotates slowly

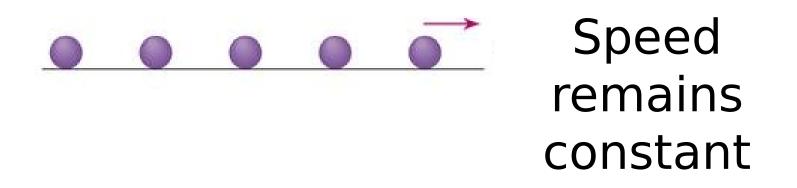


- Rings of Saturn (he couldn't see them well enough to know they were rings)
- Galileo also made essential contributions to physics, studying the nature of motion and gravity in particular

- What does an object do when there are no forces acting on it? Greeks: It comes to rest
- Is the Earth at rest?
- If, like the Greeks, you think the Earth is at rest in the center of the universe, it makes sense to say that an object is at rest.
- But the Earth is not at rest!
   And if it is not, what does it mean to say an object is "at rest"?
- What really happens? With no force on it, an object will not change its velocity: the direction and speed of its motion will be constant. This is indistinguishable from being at rest!







Galileo set up a gently sloping plank, some 12 yards in length, and made polished steel balls roll down a narrow groove cut into it. With this simple apparatus, he was able to verify his conjecture that the speed of fall increased uniformly with time – the law of uniform acceleration. It was one of the great moments in the history of science.

It became clear that the effect of force was not to produce motion, but to change motion: to produce acceleration. **A body on which no force acts moves at uniform speed.**  By rolling balls and timing their descent Galileo concluded that when no force acts on an object the object moves forever at constant speed in a straight line.

This is the final meaning of the Copernican revolution: we no longer know what absolute space means.

There is no measurement you can make to find out whether you are moving or at rest.

Galileo, initially supported by the Church, was ultimately brought for heresy in front of the Holy Office of the Inquisition in 1633. Giordano Bruno had been burned at the stake in Rome in 1600 for supporting the Copernican system, and Galileo recanted.

#### 1 Chron. 16:30 He fixed the earth firm and immovable.

#### Psalm 93:1 The world also is established, that it cannot be moved.

"Because I have been enjoined by this Holy Office to abandon the false opinion that the Sun is the center and immovable, I abjure, curse, and detest the said errors and heresies contrary to the said Holy Church."

#### Galileo Galilei, Rome 1633

Fall 1980: Pope John Paul II orders a new look at the evidence against Galileo. In 1992 a commission found Galileo Galilei had been "more perceptive" in his interpretation of the Bible than his prosecutors. The commission conceded that the Bible should be regarded as not always telling the literal truth, but sometimes as metaphor. Galileo was formally pardoned on October 31, 1992 almost 360 years after his trial.

#### Newton (1642-1727)

- Established the laws of classical mechanics.
- Invented the reflecting telescope.
- Invented calculus.



# What does an object do when there is no force on it?

After reading Galileo, Newton realized:

With no force on it, an object will not change its velocity: the direction and speed of its motion will be constant.

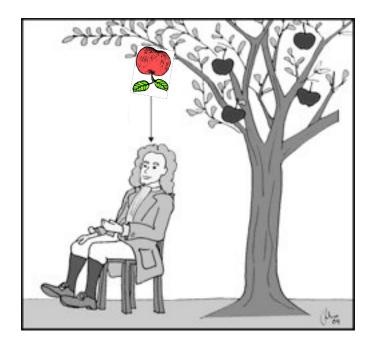
## Newton's First Law of Motion

An object at rest will stay at rest unless acted upon by a force. A moving object will continue to move forever in a straight line at constant speed, unless acted upon by a force.

> Sometimes called the **law of inertia**: **Inertia** is the tendency of an object to keep moving at the same speed and in the same direction unless acted upon by a force

Massive objects have more inertia: it takes more force to change their motion

#### Force makes an object *accelerate:* change its speed or the direction of its motion

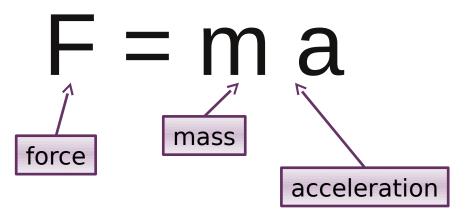


What gravity does to a falling apple is to increase the apple's speed

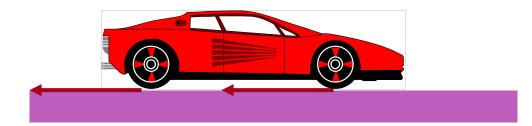
The longer an apple falls, the faster it goes

## Newton's Second Law of Motion

The acceleration of an object is directly proportional to the force applied, and inversely proportional to the object's mass.



The same force will make an object with a small mass accelerate more than an object with a larger mass.



When you slam on your brakes, the force ("friction") of the pavement on the tires of your car stops the car.

...unless the pavement is icy and the force of friction much smaller. Then the car changes its speed much more slowly.

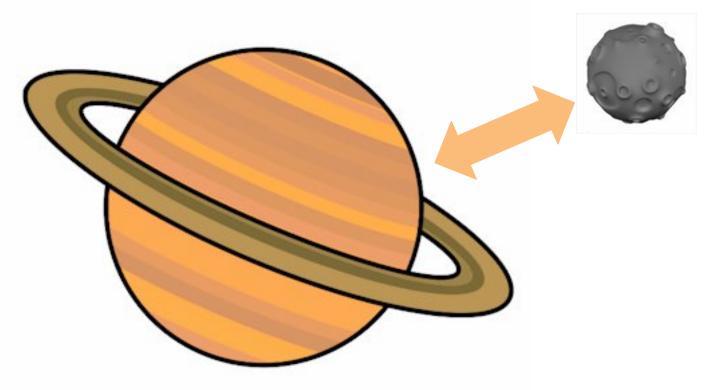
For a perfectly frictionless sheet of ice, the car would slide in a straight line at the same speed forever. Friction from the air and ground is always present under normal circumstances

This is what led the Greeks to believe that the natural tendency of objects is to come to rest



## Newton's Third Law of Motion

For every action there is an equal and opposite reaction: if body A exerts a force on body B, then body B exerts a force on body A that is equal in magnitude but oppositely directed.





The planet exerts more force on the asteroid than the asteroid exerts on the planet.

|--|

The asteroid exerts more force on the planet than the planet exerts on the asteroid.



The asteroid and the planet exert the same forces on each other.



The planet exerts more force on the asteroid than the asteroid exerts on the planet.

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The asteroid and the planet exert the same forces on each other.



The acceleration of the planet is greater than the acceleration of the asteroid.

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The acceleration of the asteroid is greater than the acceleration of the planet.



The asteroid and the planet experience the same acceleration.



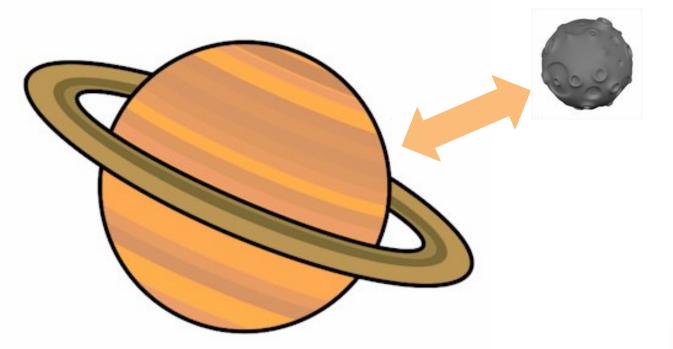
The acceleration of the planet is greater than the acceleration of the asteroid.

The acceleration of the asteroid is greater than the acceleration of the planet.



The asteroid and the planet experience the same acceleration.

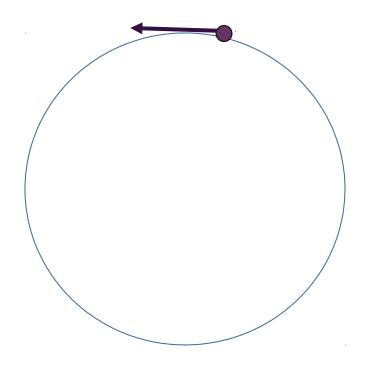
Newton's third law: the low mass asteroid and the high mass planet exert equal and opposite forces on each other when they collide.



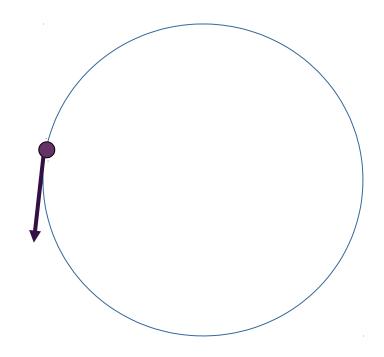
F = ma

Newton's second law: in response to the force, the low mass asteroid accelerates more than the high mass planet.

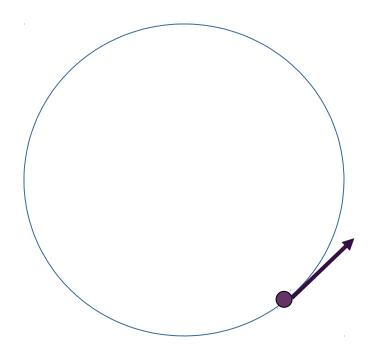
#### Is there a force on an object that moves at constant speed in a circle?



# Is there a force on an object that moves at constant speed in a circle?

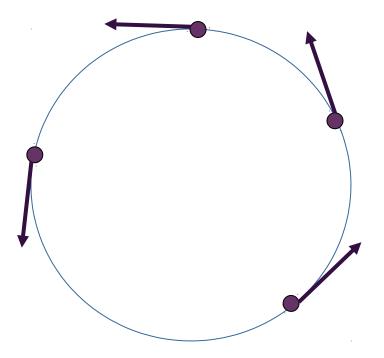


#### Is there a force on an object that moves at constant speed in a circle?



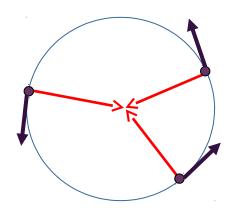
# Is there a force on an object that moves at constant speed in a circle?

Yes: The speed is constant but the **direction** of the motion changes.



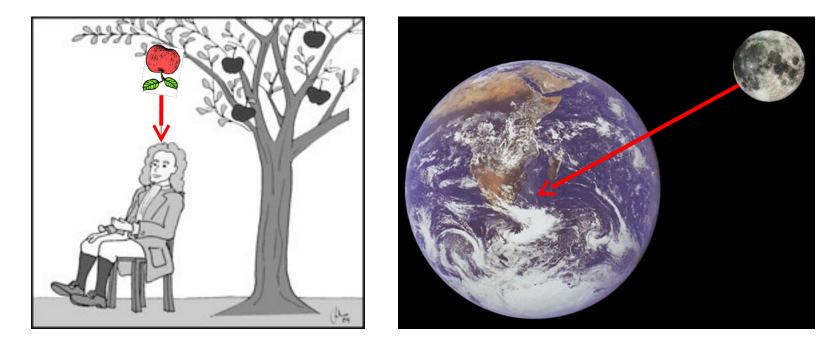
In what direction is the force on an object that moves at constant speed in a circle?

Toward the center of the circle: This is the force of a cord on a ball you swing around in a circle

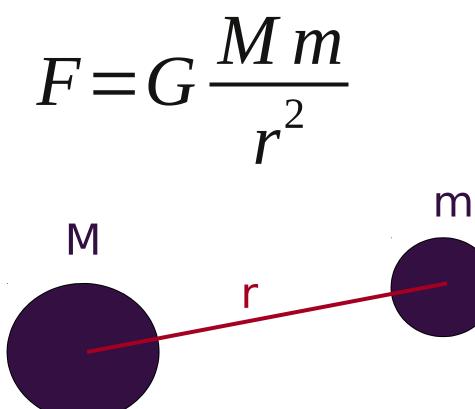




# Gravity



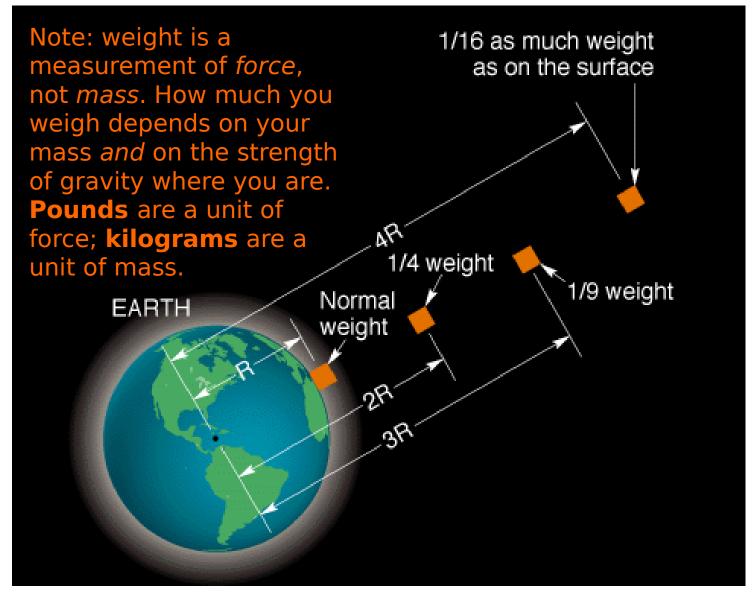
The force on the apple and the force on the Moon both point toward the center of the Earth **Newton's law of gravity**: Any two objects in the universe attract each other with a force given by



- With *r* in meters *m* in kg
- *F* in Newtons, where 1 Newton = 4.5 lb
- *G:* Newton's gravitational constant  $G = 7 \times 10^{-11}$

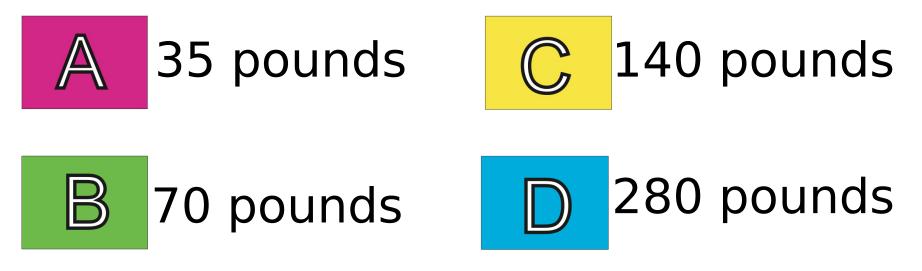
M: mass of object 1 m: mass of object 2 r: distance between the objects

# Force of gravity on an object proportional to 1/r<sup>2</sup>



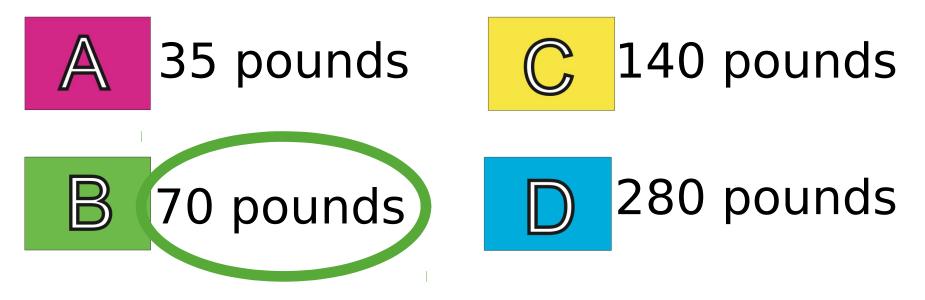
# Using Newton's law of Gravity $F = G \frac{Mm}{r^2}$

An astronaut weighs 140 pounds on Earth. How much would she weigh on a planet that is the same size as Earth but ½ the mass?



# Using Newton's law of Gravity $F = G \frac{Mm}{r^2}$

An astronaut weighs 140 pounds on Earth. How much would she weigh on a planet that is the same size as Earth but ½ the mass?



$$F = G \frac{Mm}{r^2}$$
Mass of new planet is ½ mass of Earth:  

$$F_{new} = G \frac{(0.5M)m}{r^2} = 0.5 G \frac{Mm}{r^2} = 0.5 F_{old}$$

So the gravitational force on the new planet is half as strong as on Earth, and the astronaut weighs half as much.

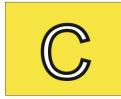
Remember that weight is gravitational force! The **mass** of the astronaut *m* does not change!

## Using Newton's law of Gravity $F = G \frac{Mm}{r^2}$

An astronaut weighs 150 pounds on Earth. How much would she weigh on a planet that is the same size as Earth but 0.1 times the mass?







150 pounds



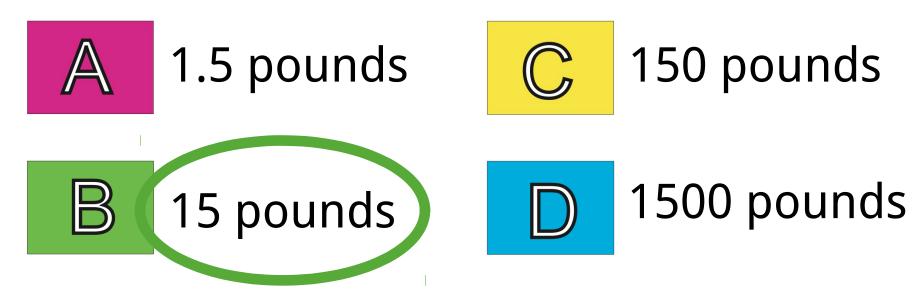




1500 pounds

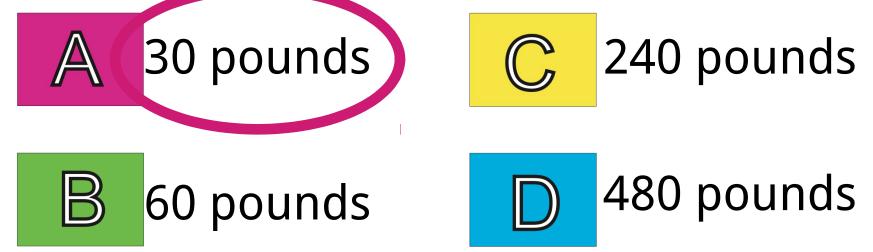
### Using Newton's law of Gravity $F = G \frac{Mm}{r^2}$

An astronaut weighs 150 pounds on Earth. How much would she weigh on a planet that is the same size as Earth but 0.1 times the mass?



### Using Newton's law of Gravity $F = G \frac{Mm}{r^2}$

An astronaut weighs 120 pounds on Earth. How much would she weigh on a planet that is the same mass as Earth but double the radius?



$$F = G \frac{Mm}{r^2}$$

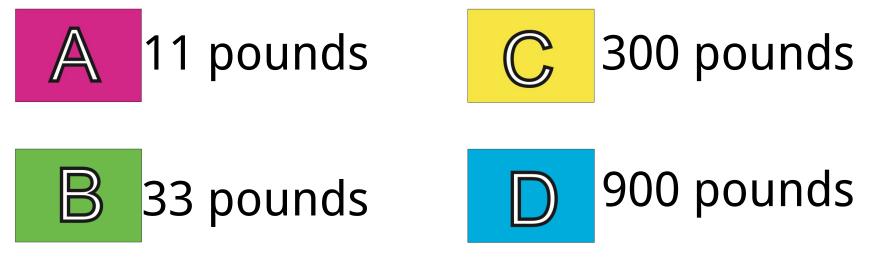
Radius of new planet is twice the radius of Earth:

$$F_{new} = G \frac{Mm}{(2r)^2} = \frac{1}{2^2} G \frac{Mm}{r^2} = \frac{1}{4} F$$

The gravitational force on the new planet is ¼ as strong as on Earth, and the astronaut weighs ¼ as much.

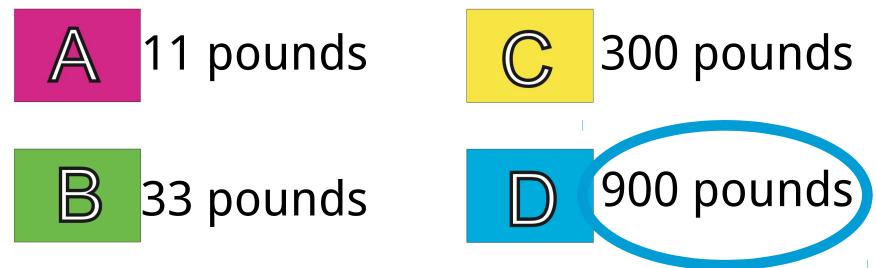
# Using Newton's law of Gravity $F = G \frac{Mm}{r^2}$

An astronaut weighs 100 pounds on Earth. How much would she weigh on a planet that is the same mass as Earth but 1/3 the radius?



# Using Newton's law of Gravity $F = G \frac{Mm}{r^2}$

An astronaut weighs 100 pounds on Earth. How much would she weigh on a planet that is the same mass as Earth but 1/3 the radius?



$$F = G \frac{Mm}{r^2}$$
Radius of new planet is 1/3 the radius of Earth:

$$F_{new} = G \frac{Mm}{\left(\frac{1}{3}r\right)^2} = \frac{1}{\left(\frac{1}{3^2}\right)} G \frac{Mm}{r^2} = 9F$$

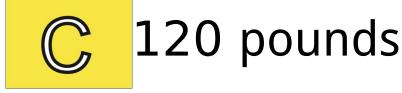
The gravitational force on the new planet is 9 times stronger than it is on Earth, and the astronaut weighs 9 times as much!

## Using Newton's law of Gravity $F = G \frac{Mm}{r^2}$

An astronaut weighs 120 pounds on earth. How much would she weigh on a planet that is both  $\frac{1}{2}$  the mass and  $\frac{1}{2}$  the radius of Earth?







240 pounds

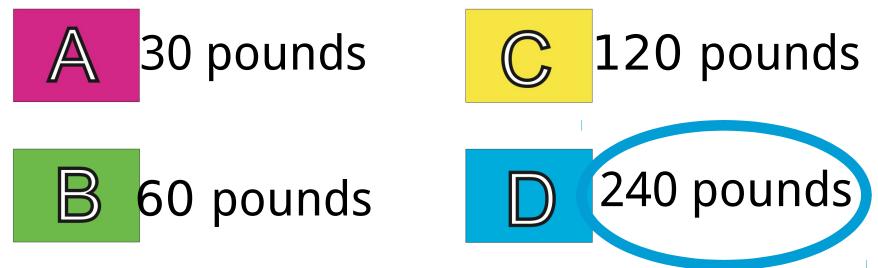




Using Newton's law of Gravity  

$$F_{new} = G \frac{(0.5M)m}{(0.5r)^2} = \frac{0.5}{0.25} G \frac{Mm}{r^2} = 2F$$

An astronaut weighs 120 pounds on earth. How much would she weigh on a planet that is both  $\frac{1}{2}$  the mass and  $\frac{1}{2}$  the radius of Earth?



### Using Newton's law of Gravity $F = G \frac{Mm}{r^2}$

The International Space Station orbits at about 400 km above the Earth's surface, or about 6770 km from the center of the Earth. The radius of the Earth is about 6370 km. Which of the following is true?



The force of gravity on the space station is much weaker than it is on the surface of the Earth.

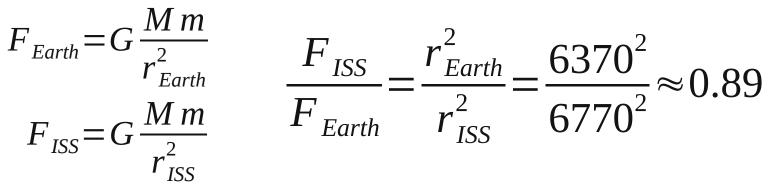


The force of gravity is almost as strong on the space station as it is on the surface of the Earth.



There is no force of gravity on the space station.

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## So why are astronauts on the Space Station weightless??

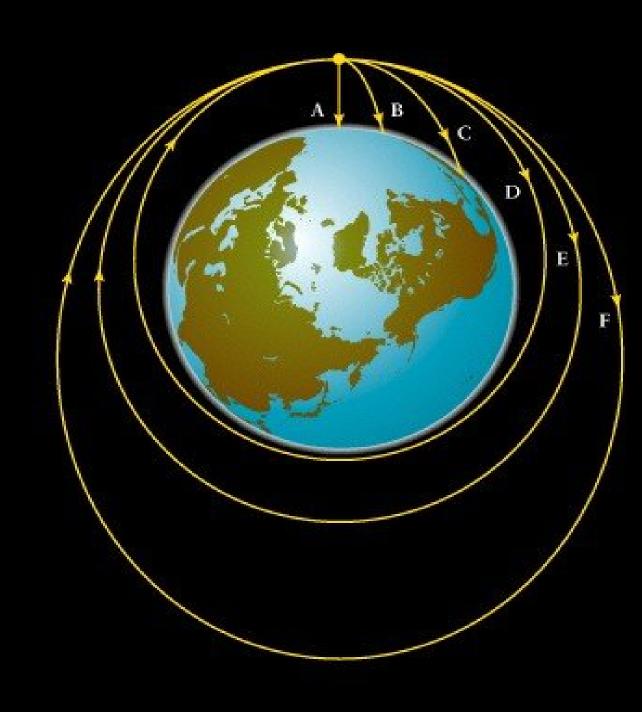


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#### So why are astronauts on the Space Station weightless??

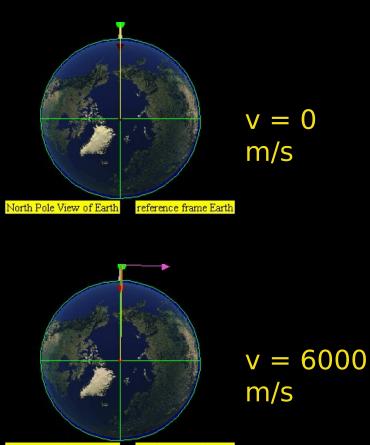
- The space station orbits due to Earth's gravity – if this disappeared it would fly off into space in a straight line
- It does not fall back to Earth because it's going very fast – in fact it's always falling, but it falls "over the horizon"
- The astronauts do not feel the force of gravity because they are in free fall





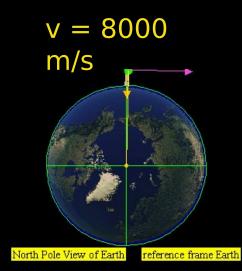
"Newton's Cannon"

A cannonball launched from above the Earth's surface will, if launched fast enough, orbit the Earth instead of falling back to the ground.

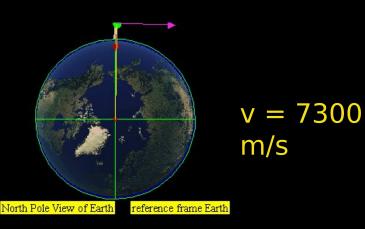


North Pole View of Earth

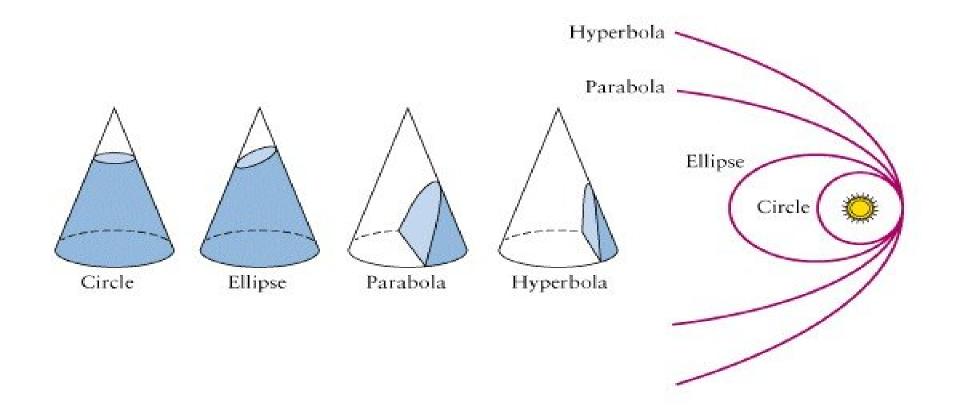
reference frame Earth



v = 10,000 m/s



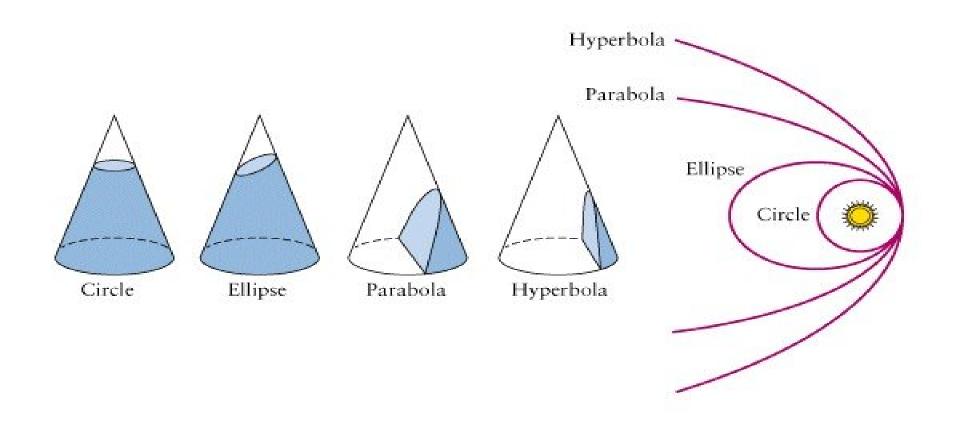
There are two types of orbits, depending on whether an object is moving faster than its escape velocity. Bound orbits are **ellipses** Unbound orbits are **hyperbolas** 



Turns out these shapes can be constructed by taking so-called conic sections

#### Just to reiterate:

#### *Bound orbits are ellipses Unbound orbits are* **hyperbolas**



## Newton's Theory of Gravity provided an explanation for Kepler's Laws

• Kepler's second law:

## Newton's Theory of Gravity provided an explanation for Kepler's Laws

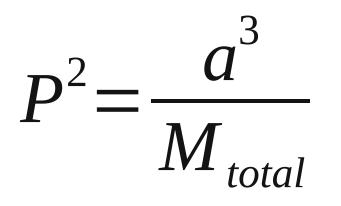
 Kepler's second law: The line from the Sun to a planet sweeps out equal areas in equal intervals of time. This means the planet moves fastest when it's closest to the Sun (perihelion) and slowest when it's farthest from the Sun (aphelion).

$$F = G \frac{Mm}{r^2}$$

Newton's law of gravity tells us that the gravitational force on the planet is strongest when it's closest to the Sun: when r is smallest. Newton's second law (F = ma) tells us that force causes acceleration: as the force increases, the acceleration of the planet increases. This is why the planet moves faster as it gets closer to the Sun and slower as it gets farther away.

## Newton's Theory of Gravity provided an explanation for Kepler's Laws

- And resulted in a couple of corrections
  - Kepler's first law: The orbit of a planet around the Sun is an ellipse with the center of mass of the planet-Sun system at one focus
  - Kepler's third law: The relationship between the period and semi-major axis of an orbiting planet is



*P* in Earth years *a* in AU *M*total is the **combined mass** of the two objects *in units of the mass of the Sun* 

This correction doesn't matter much for a planet which is much, much less massive than the star it orbits, but for two objects of similar masses, like a binary star system, it is very important!

#### Predictions from Newton's Theory of Gravity

- Precise orbits of planets, moons, comets . . .
- Reappearance of Halley's comet
- Discovery of Neptune from discrepancy in orbit of Uranus. Predicted by Adams and Le Verrier

#### Reading assignment

Approximate reading schedule for the upcoming week

- Wednesday, Sept 17: Chapter 2, sections 2.1 to 2.3
- Friday, Sept 19: Chapter 2, sections 2.4 to 2.8
- Monday, Sept 22: Planetarium
- Wednesday Sept 24: Chapter 3, sections 3.1 to 3.3
- Friday Sept 26: Chapter 3, sections 3.4 and 3.5